

Jet structure modelling with self-consistent binary neutron star merger environments

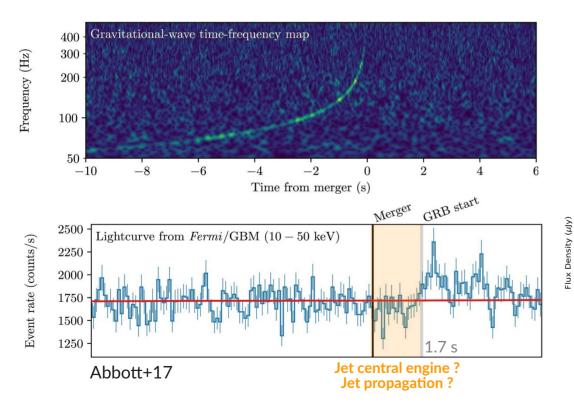
A. Pavan, R. Ciolfi, J. V. Kalinani, and A. Mignone

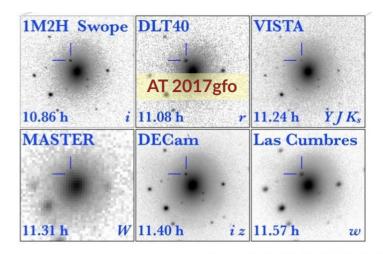


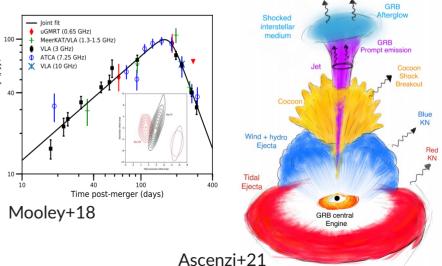




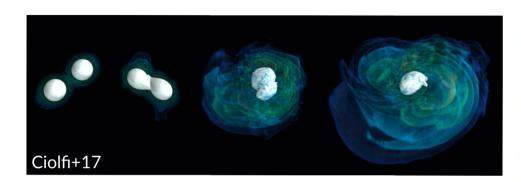
GW170817 / GRB 170817A



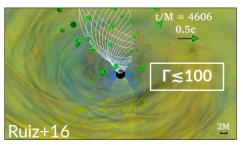




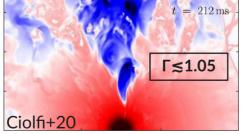
BNS Merger Simulations



BH - Accretion disk

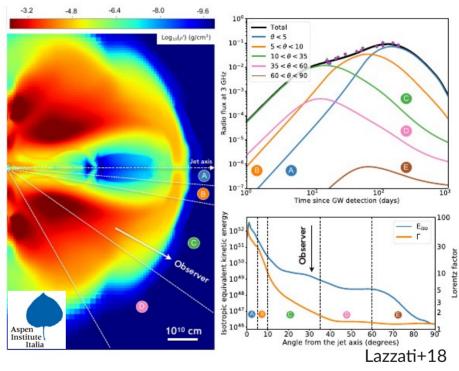


Magnetar



+ Neutrino leakage and Realistic EOS e.g., Mosta+20, Ruiz+21, Sun+22

Short-GRB Jet Simulations

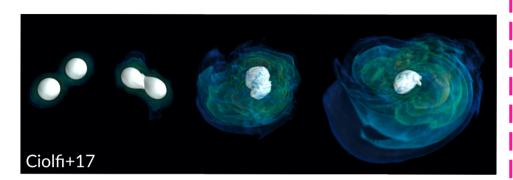


+ 3D (G)RMHD

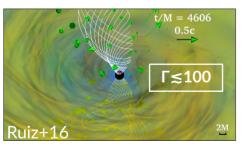
e.g., Geng+19, Nathanail+21, Gottlieb+22

NO DIRECT CONNECTION

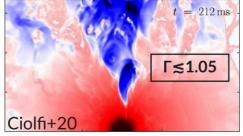
BNS Merger Simulations



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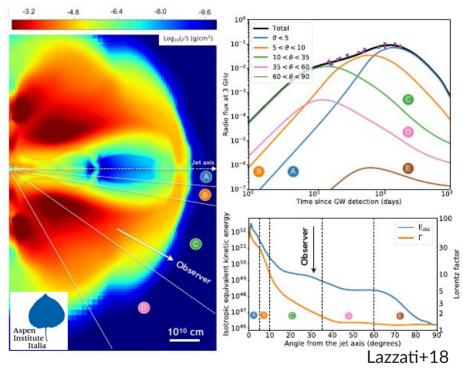


Magnetar



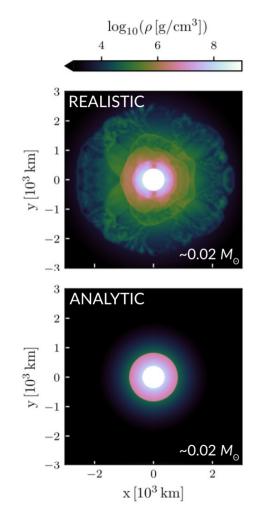
- X High computational costs
- Resolution issues (e.g., magnetic field instabilities)
- x Limited scales (≤250ms, ≤10³ km)

Short-GRB Jet Simulations



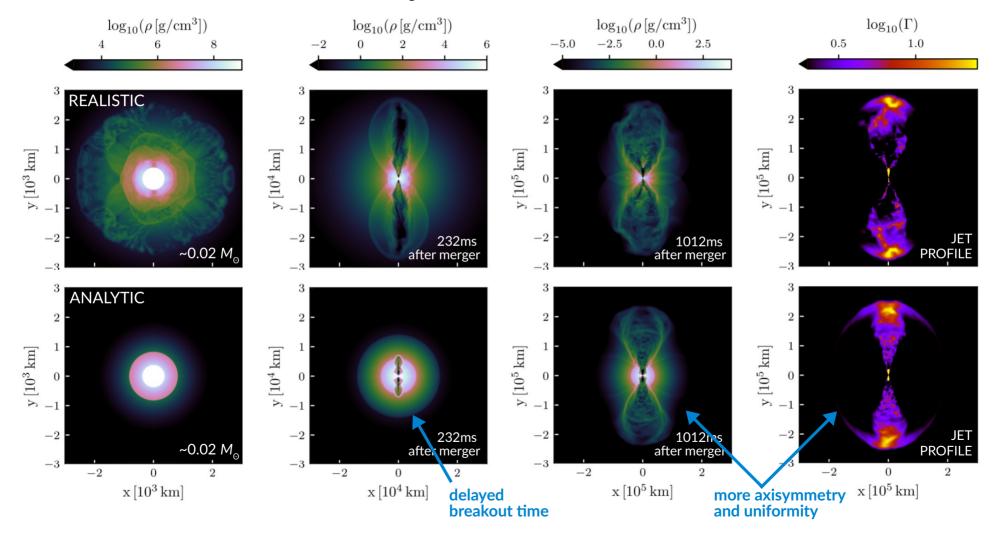
- × 3D-(G)RMHD up to ~2s, $\geq 10^5$ km
- Resolution issues (e.g., magnetic reconnection)
- X Analytic environment

Pavan+21 (RHD jet in realistic environment)



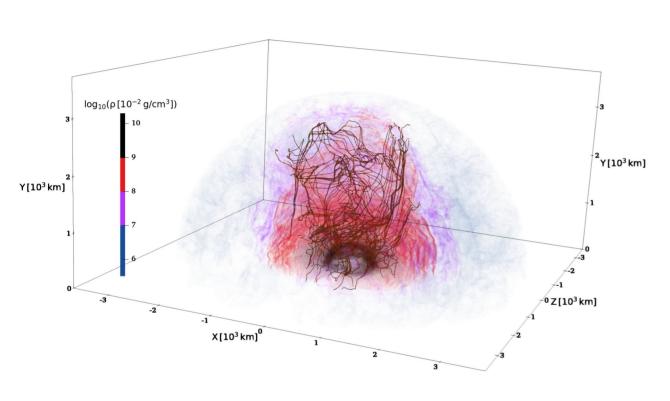
- 3D-GRHD BNS merger simulation by Kalinani (WhiskyMHD, Einstein Toolkit, PostCactus)
- Data import into special-relativistic code PLUTO (Mignone+07,+12)
- Domain setting
 - 3D spherical coordinates
 - Logarithmic radial spacing + "excision"
 - Power-law "atmosphere" (static)
- Newtonian gravity included
- Jet injection environment
 - REALISTIC \rightarrow Self-consistent [P,p,v](r, θ , ϕ) and boundary conditions
 - ANALYTIC → Power-law + Homologous expansion

Pavan+21 (RHD jet in realistic environment)





Pavan+22 (RMHD jet in realistic environment)



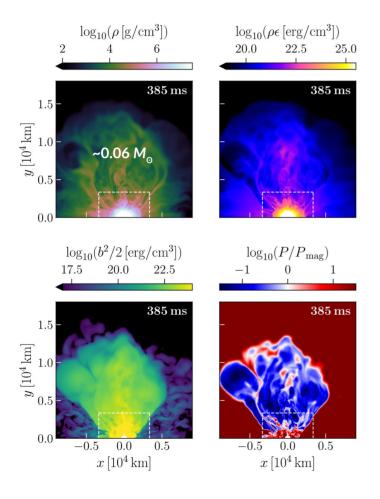
 $\log_{10}(\rho \,[\mathrm{g/cm^3}])$ 5 $155\,\mathrm{ms}$ BNS $y \left[10^3 \, \mathrm{km} \right]$ $log_{10}(B[G])$ 13 $155\,\mathrm{ms}$ $y [10^3 \, \mathrm{km}]$ $x \, [10^3 \, \mathrm{km}]$

3D-GRMHD BNS merger simulation by Ciolfi20 (WhiskyMHD, Einstein Toolkit, PostCactus)

Import into PLUTO-RMHD (Hyperbolic Divergence Cleaning)



Pavan+22 (RMHD jet in realistic environment)



Jet injection setting

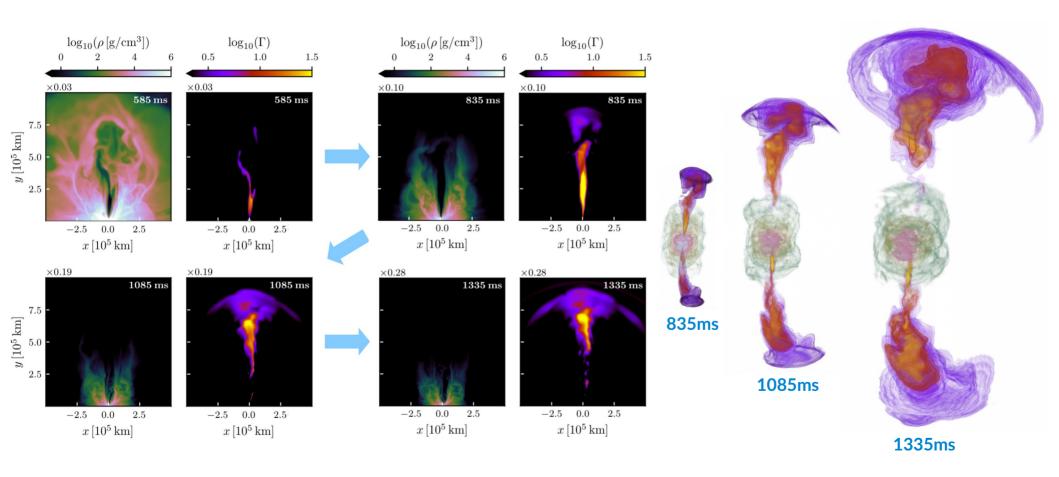
- 385ms after merger
- Newtonian gravity included
- Semi-analytic jet (Martí15, Geng+19)
 - Transversal equilibrium
 - Uniform rotation
 - Decaying luminosity ($\tau = 0.3s$)

Table 1. Jet Injection Parameters

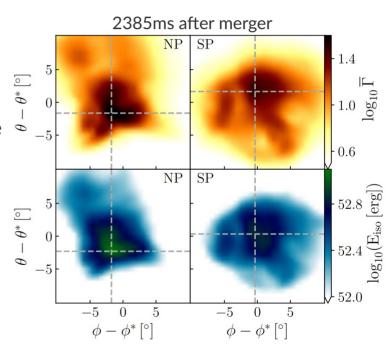
$\Delta t_{\rm j}$ [ms]	$L_{ m j}$ [erg/s]	3	$\Gamma_{ m j}$	Γ_{∞}	θ _{j,m} [°]	B_{j}^{r} [G]	$B_{\mathrm{j,m}}^{\phi}$ [G]
385	$3 \cdot 10^{51}$	10	3	300	4	$3.1 \cdot 10^{12}$	$6.2 \cdot 10^{12}$



Pavan+22 (RMHD jet in realistic environment)

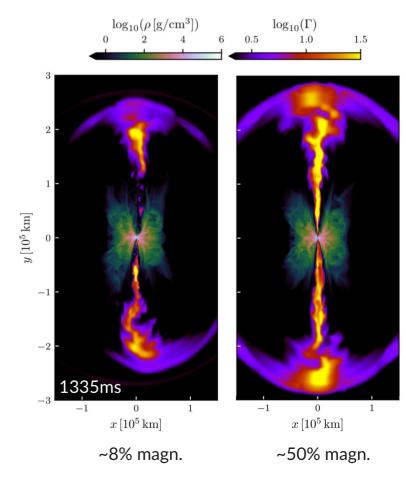


- ✓ SGRB jet simulations in fully-realistic BNS merger environments
- Realistic environments play a crucial role
 (jet breakout time, final structure & energetic)
- ✓ Jet injection time (with respect to merger) has strong influence
- Magnetic fields & Newtonian gravity must be included
- Dependence on jet injection parameters
- X Afterglow modelling
- x Realistic jet imported from BNS merger simulations

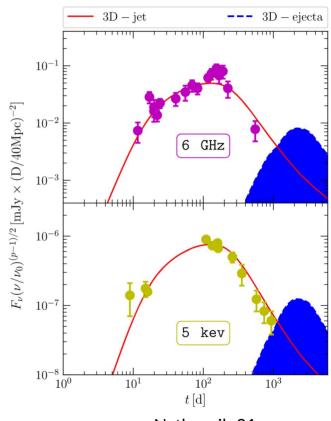


Structured jets with 3D-asymmetries

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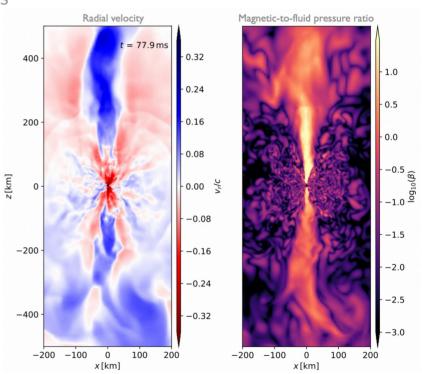
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Nathanail+21

✓ SGRB jet simulations in fully-realistic BNS merger environments

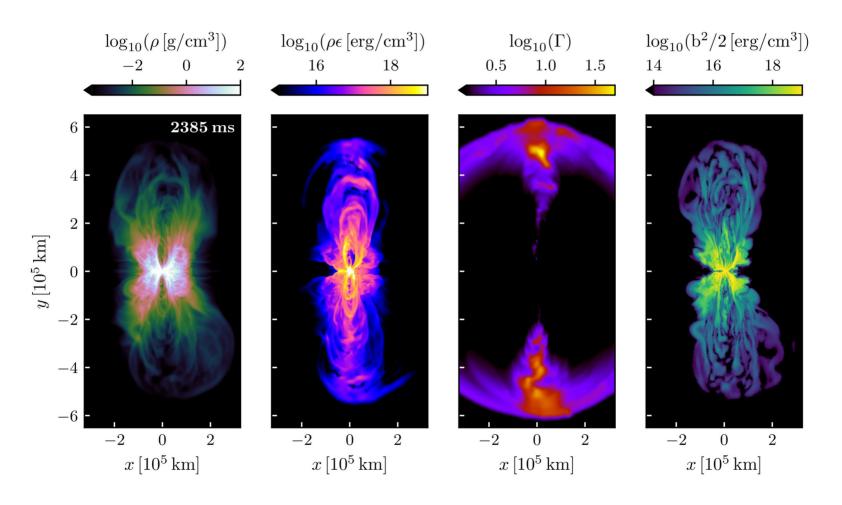
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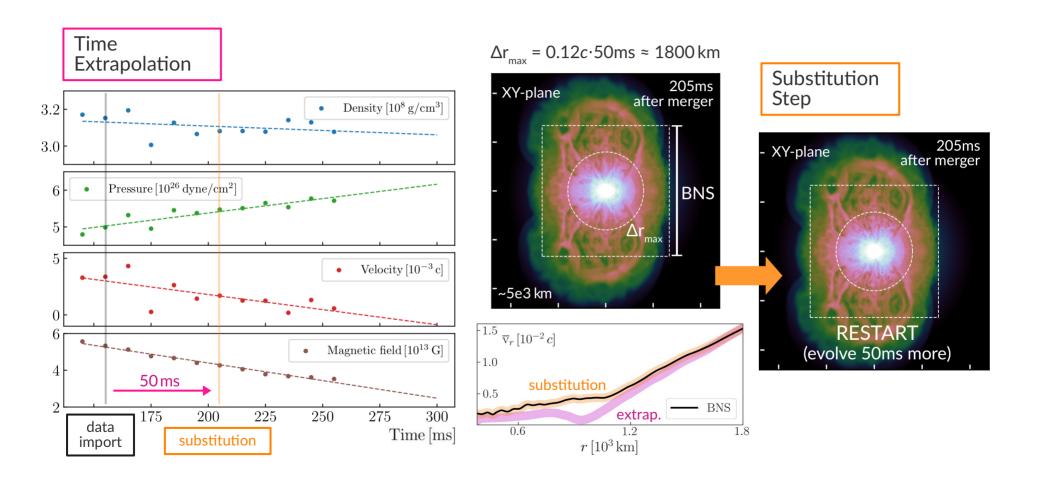
Simulation by Kalinani



Configuration at ~2s



Pavan+22 (extrapolation procedure)



Newtonian gravity impact

